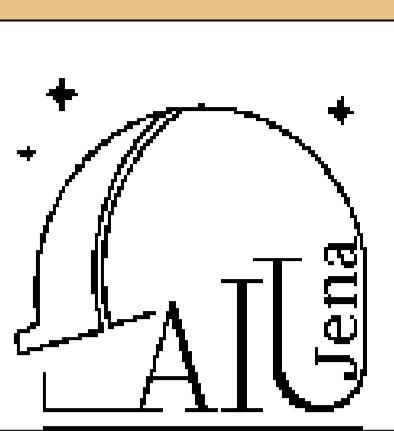


Semiconductor- & Metallic-Behavior of Iron(II) Sulfides: A Multi-Wavelengths Study

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Abstract

- Iron(II) sulfides: **Sulfur reservoirs in interplanetary dust** (e.g. in comets, meteorites, IDPs)
- Prior spectral studies: Covers either the UV-VIS to NIR, or the FIR spectral range
- Our study: Reflectivity data from **VUV to FIR** for the determination of optical constants
- Five iron(II) sulfides with different Fe/S ratio: From **4C-pyrrhotite to troilite**
- Binary spectroscopic behavior** between stoichiometric(-like) & iron depleted iron(II) sulfides
- Optical constants** depend on Fe/S ratio & **IR reflectivity decreases with higher iron amount**

Previous Optical Studies

- [1] W. G. Egan, T. Hilgeman, The interstellar medium: UV complex index of refraction of several candidate materials, *Astronomical Journal* 80, 1975, pp. 587-594.
- [2] W. G. Egan, T. Hilgeman, The Rings of Saturn: A Frost-Coated Semiconductor?, *Icarus* 30(2) 1977, pp. 413-421.
- [3] B. Begemann, J. Dorschner, T. Henning, H. Mutschke, E. Thamm, A Laboratory Approach to the Interstellar Sulfide Dust Problem, *E. 1994, ApJ, 423, L71-L74.*
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- [5] J. B. Pollack, D. Hollenbach, S. Beckwith, D. P. Simonelli, T. Roush, W. Fong, Composition and Radiative Properties of Grains in Molecular Clouds and Accretion Disks, *The Astrophysical Journal* 421 1994, pp. 615-639.
- [6] T. Henning, H. Mutschke, Low-temperature infrared properties of cosmic dust analogues., *A&A*, 1997, pp. 327-743.

Samples

	Troilite	Synthetic [3,4,6]	DAL [7]	NYS [8]	TYS [8]	DEV [9]
Iron(II) Sulfide Type	Troilite	Troilite with iron	4C-Pyrrhotite	NC-Pyrrhotite: 4.78 – 4.96	NC-Pyrrhotite: 5.12 – 5.52	6C-Pyrrhotite
Iron/Sulfur Ratio	1	1	0.877(1)	0.896(2)	0.903(7)	0.929(2)
Crystal System	Hexagonal	Hexagonal	Monoclinic	Hexagonal		Monoclinic
Origin	MPIK Heidelberg	Begemann, MPS Jena	Harries, IGW Jena		de Velliers	
Images						

Experimental Setup

- Four spectrometers from **VUV to mm wavelength range at 300 K**
- Two **frequency domain** spectrometers:
 - VUV: 115 nm – 230 nm (under vacuum)
 - UV-VIS-NIR: 190 nm – 2.5 μm (under ambient air)
- Two **time domain** spectrometers:
 - FTIR: 2.0 μm – 500 μm (under vacuum)
 - THz: 250 μm – 2 mm (under dry air)
- Low temperature measurements down to 10 K for MIR & FIR**

Multi-Wavelengths Reflectivity

- Binary spectroscopic behavior:**
 - Semiconductor-like** for stoichiometric troilite
 - Metal-like** for iron deficient pyrrhotites
- Mid-IR reflectivity deceases with increasing Fe/S ratio (@ 10 μm)

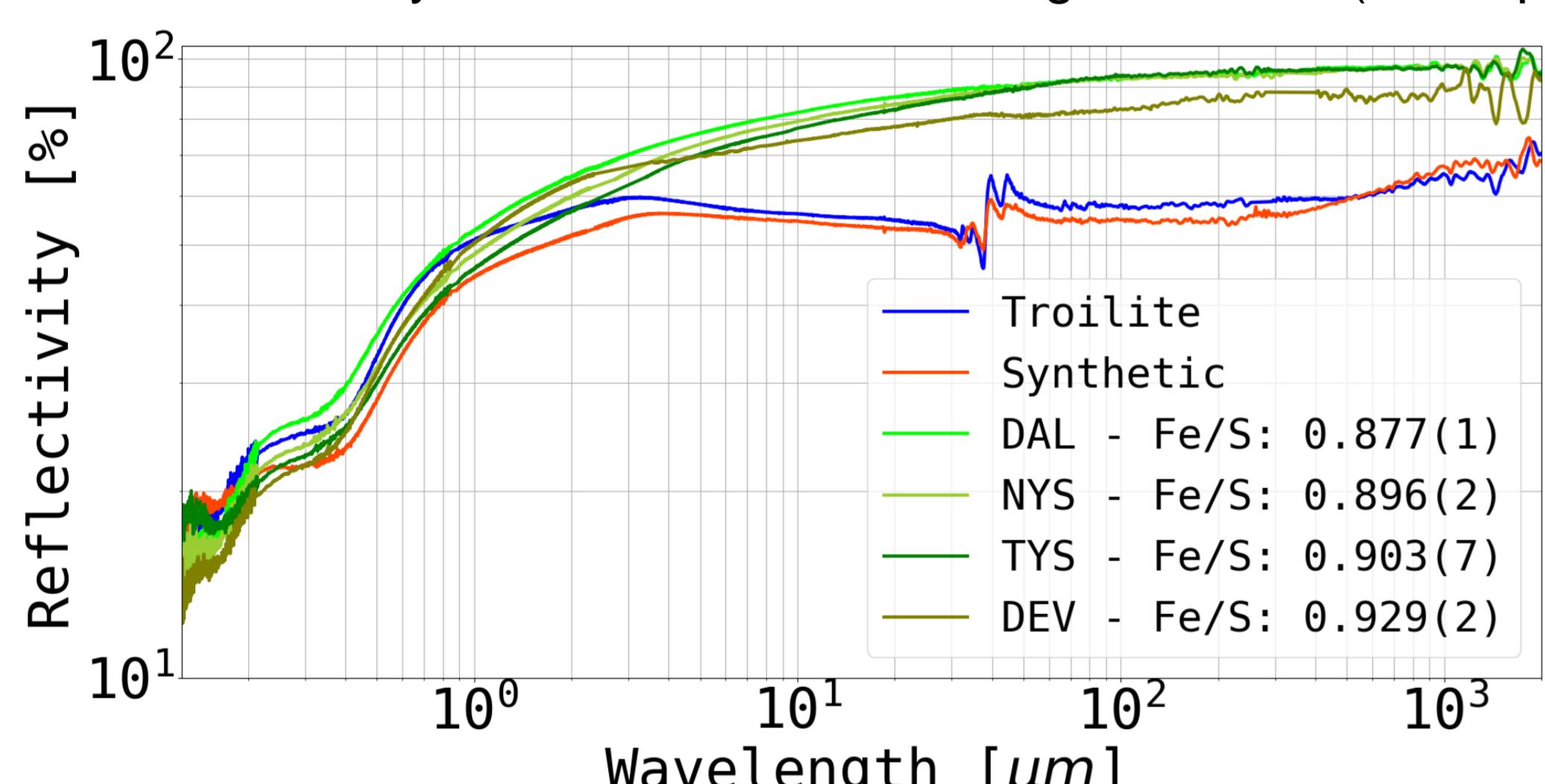


Figure 1: Room temperature reflectivity for samples troilite and pyrrhotites from VUV to mm wavelength regime.

Optical Constants

- Fitting routine** based on **Levenberg-Marquardt** method
- Optimize n, k parameters based on comparison between measured and computed reflectivity
- Drude-Lorentz-oscillator-model** for troilite & pyrrhotites
- New data in agreement with measured literature values

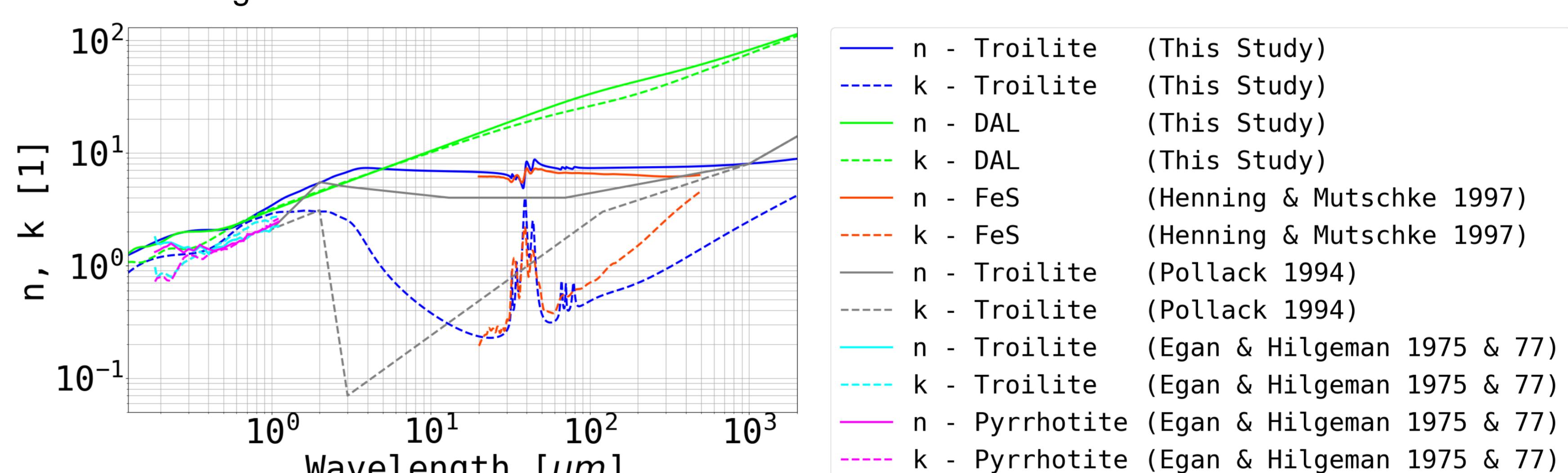


Figure 2: Comparison of n & k values for samples troilite and 4C-pyrrhotite with literature [1, 2, 5, 6].

Low Temperature Measurements

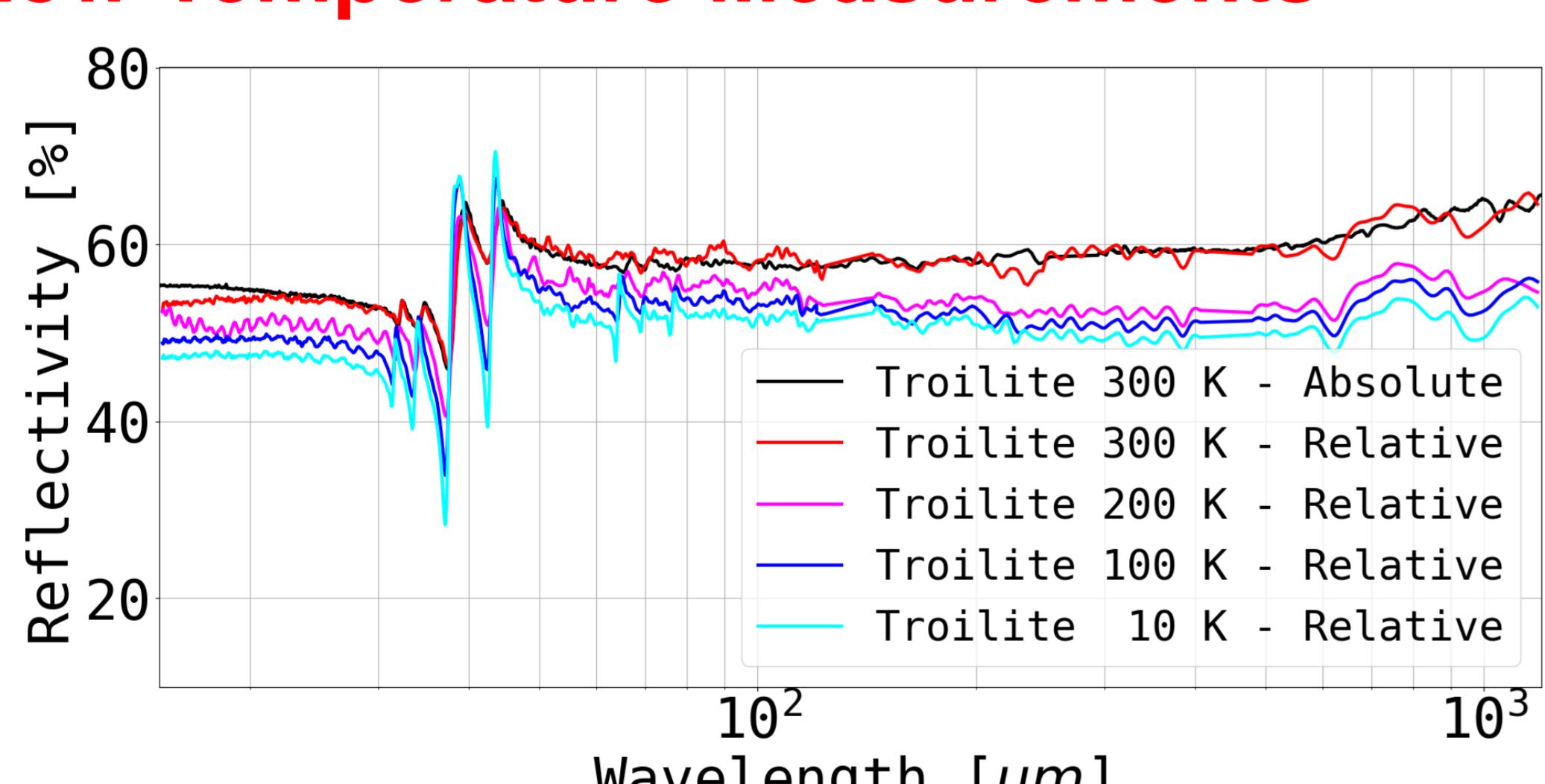


Figure 3: MIR and FIR reflectivity of troilite at low temperatures.

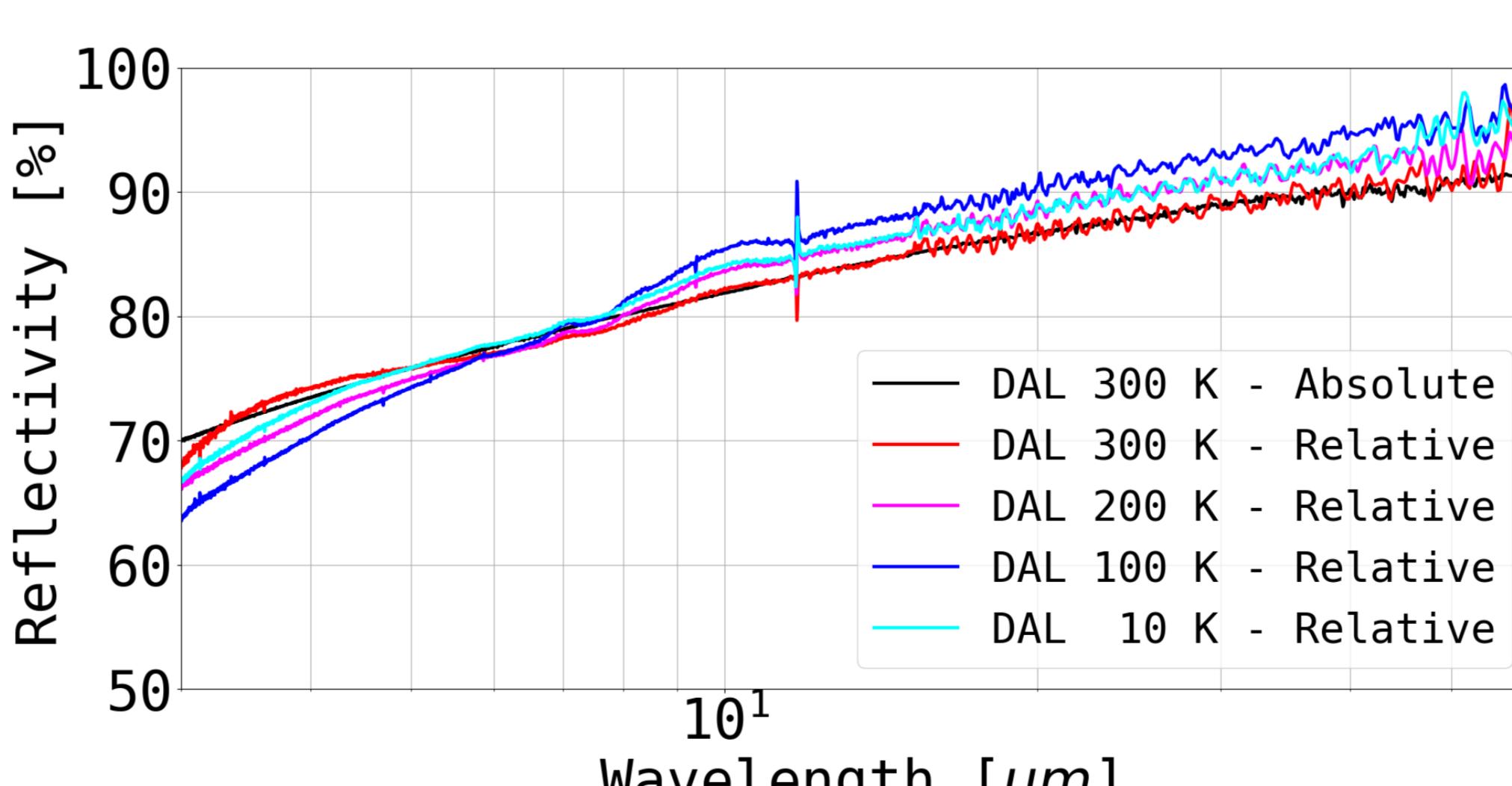


Figure 4: MIR reflectivity of 4C-pyrrhotite at low temperatures.

- Troilite – Semiconductor:**
 - Seven phonon bands:** 32.4 μm, 34.9 μm, 39.7 μm, 44.5 μm, 67.1 μm, 71.1 μm, 79.7 μm (at room temperature)
 - Blue-shift & sharpening of phonon bands
- 4C-Pyrrhotite – Metal:**
 - Free charge carrier characteristics**
 - Change of sign for reflectivity change in IR
 - Possible origin: **Besnus transition at 32 K** [10]

Acknowledgements

This work was conducted in the Research Unit FOR 2285 "Debris Disks in Planetary Systems" of the Deutsche Forschungsgemeinschaft (grant MU1164/9-2).

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